Available online at www.pelagiaresearchlibrary.com



Pelagia Research Library

Advances in Applied Science Research, 2012, 3 (4):2312-2325



Influence of heat transfer on MHD oscillatory flow of Jeffrey fluid in a channel

K. Kavita¹, K. Ramakrishna Prasad² and B. Aruna Kumari³

¹Department of Mathematics, Tirumala Engineering College, Bogaram, Hyderabad, A.P., INDIA ²Department of Mathematics, S V University, Tirupati, A.P, INDIA ³Chadalawada Engineering College, Titupathi, A.P, INDIA

ABSTRACT

In this paper, Influence of heat transfer on MHD oscillatory flow of Jeffrey fluid in a channel is investigated. The effects of various emerging parameters on the velocity field and temperature field are discussed in detail through graphs.

Keywords: Jeffrey Fluid, MHD, Oscillatory flow

INTRODUCTION

The flow of an electrically conducting fluid has important applications in many branches of engineering science such as magnetohydrodynamics (MHD) generators, plasma studies, nuclear reactor, geothermal energy extraction, electromagnetic propulsion, the boundary layer control in the field of aerodynamics and so on. Heat transfer effect on laminar flow between parallel plates under the action of transverse magnetic field was studied by Nigam and Singh (1960). Soundalgekar and Bhat (1971) have investigated the MHD oscillatory flow of a Newtonian fluid in a channel with heat transfer. MHD flow of viscous fluid between two parallel plates with heat transfer was discussed by Attia, and Kotb (1996). Raptis et al. (1982) have analyzed the hydromagnetic free convection flow through a porous medium between two parallel plates. Aldoss et al. (1995) have studied mixed convection flow from a vertical plate embedded in a porous medium in the presence of a magnetic field. Makinde and Mhone (2005) have considered heat transfer to MHD oscillatory flow in a channel filled with porous medium. Mostafa (2009) have studied thermal radiation effect on unsteady MHD free convection flow past a vertical plate with temperature dependent viscosity. Unsteady heat transfer to MHD oscillatory flow through a porous medium under slip condition was investigated by Hamza et al. (2011).

Moreover the non-Newtonian fluids are more appropriate than Newtonian fluids in many practical applications. Examples of such fluids include certain oils, lubricants, mud, shampoo, ketchup, blood at low shear rate, cosmetic products, polymers and many others. Unlike the viscous fluids, all the non-Newtonian fluids (in terms of their diverse characteristics) cannot be described by a single constitutive relationship. Hence, several models of non-Newtonian fluids are proposed in the literature. Al Khatib andWilson (2001) have studied the Poiseuille flow of a yield stress fluid in a channel. Flow of a visco-plastic fluid in a channel of slowly varying width was studied by Frigaard and Ryan (2004). Aamir Ali and Saleem Asghar (2011) have analyzed by oscillatory channel flow for non-Newtonian fluid. Rita and Jyoti Das (2012) have studied the effect of heat transfer on MHD oscillatory viscoelastic fluid flow in a channel through a porous medium.

Pelagia Research Library

K. Kavita et al

In view of these we studied the effect of heat transfer on MHD oscillatory flow of a Jeffrey fluid in a channel. The expressions are obtained for velocity and temperature analytically. The effects of various emerging parameters on the velocity and temperature are discussed through graphs in detail.

2. Mathematical Formulation

We consider the flow of a Jeffrey fluid in a channel of width h under the influence of electrically applied magnetic field and radiative heat transfer as depicted in Fig.1. It is assumed that the fluid has small electrical conductivity and the electromagnetic force produced is very small. We choose the Cartesian coordinate system (x, y), where x- is taken along center of the channel and the y- axis is taken normal to the flow direction.



Fig. 1 Physical model of the problem

The constitute equation of S for Jeffrey fluid is

$$S = \frac{\mu}{1 + \lambda_1} \left(\dot{\gamma} + \lambda_2 \ddot{\gamma} \right) \tag{2.1}$$

where μ is the dynamic viscosity, λ_1 is the ratio of relaxation to retardation times, λ_2 is the retardation time, $\dot{\gamma}$ is the shear rate and dots over the quantities denote differentiation with time.

The basic equations of momentum and energy governing such a flow, subject to the Boussinesq approximation, are:

$$\rho \frac{\partial u}{\partial t} = -\frac{\partial p}{\partial x} + \frac{\mu}{1 + \lambda_1} \frac{\partial^2 u}{\partial y^2} - \sigma B_0^2 u + \rho g \beta (T - T_0)$$
(2.2)

$$\rho \frac{\partial T}{\partial t} = \frac{k}{c_p} \frac{\partial^2 T}{\partial y^2} - \frac{1}{c_p} \frac{\partial q}{\partial y}$$
(2.3)

The boundary conditions are given by

$$u = 0, T = T_0 at y = 0$$
 (2.4)

$$u = 0, \qquad T = T_1 \qquad \text{at} \qquad y = h \tag{2.5}$$

where u is the axial velocity, T is the fluid temperature, p is the pressure, ρ is the fluid density, B_0 is the magnetic field strength, σ is the conductivity of the fluid, g is the acceleration due to gravity, β is the coefficient of volume expansion due to temperature, c_p is the specific heat at constant pressure, k is the thermal

Pelagia Research Library

2313

conductivity and q is the radiative heat flux. Following Cogley et al. (1968), it is assumed that the fluid is optically thin with a relatively low density and the radiative heat flux is given by

$$\frac{\partial q}{\partial y} = 4\alpha^2 \left(T_0 - T\right) \tag{2.6}$$

here α is the mean radiation absorption coefficient.

Introducing the following non-dimensional variables

$$\bar{x} = \frac{x}{h}, \ \bar{y} = \frac{y}{h}, \ \bar{u} = \frac{u}{U}, \ \theta = \frac{T - T_0}{T_1 - T_0}, \ \bar{t} = \frac{tU}{h}, \ \bar{p} = \frac{pa}{\mu U}, \ M^2 = \frac{\sigma h^2 B_0^2}{\mu}, \ Gr = \frac{\rho g \beta (T_1 - T_0)}{U\mu}, \ Re = \frac{\rho h U}{\mu}, \ Pe = \frac{\rho h Uc_p}{k}, \ N^2 = \frac{4\alpha^2 h^2}{k}$$

here U is the mean flow velocity, into the equations (2.2) and (2.3), we get (after dropping bars)

$$\operatorname{Re}\frac{\partial u}{\partial t} = -\frac{\partial p}{\partial x} + \frac{1}{1+\lambda_1}\frac{\partial^2 u}{\partial y^2} - M^2 u + Gr\theta$$
(2.7)

$$Pe\frac{\partial T}{\partial t} = \frac{\partial^2 T}{\partial y^2} + N^2\theta$$
(2.8)

where Re is the Reynolds number, M is the Hartmann number, Gr is the Grashof number, Pe is the Peclet number and N is the radiation parameter.

The corresponding non-dimensional boundary conditions are

$$u = 0, \qquad \theta = 0 \qquad \text{at} \qquad y = 0$$
 (2.9)
 $u = 0, \qquad \theta = 1 \qquad \text{at} \qquad y = 1$ (2.10)

3. Solution

12

In order to solve equations (2.7) - (2.10) for purely oscillatory flow, let

$$-\frac{\partial p}{\partial x} = \lambda e^{i\omega t}$$
(3.1)

$$u(y,t) = u_0(y)e^{i\omega t}$$

$$(3.2)$$

$$\theta(y,t) = \theta_0(y)e^{i\omega t}$$
(3.3)

where λ is a real constant and ω is the frequency of the oscillation.

Substituting the equations (3.1) - (3.3) in to the equations (2.7) - (2.10), we get

$$\frac{d^{2}u_{0}}{dy^{2}} - m_{2}^{2}u_{0} = -\lambda(1+\lambda_{1}) - Gr(1+\lambda_{1})\theta_{0}$$
(3.4)

$$\frac{d^2\theta_0}{dy^2} + m_1^2\theta_0 = 0$$
(3.5)

with the boundary conditions

$$u_0 = 0, \qquad \theta_0 = 0 \qquad \text{at} \qquad y = 0$$
 (3.6)

Pelagia Research Library

2314

$$u_0 = 0, \quad \theta_0 = 1 \quad \text{at} \quad y = 1$$

in which $m_1 = \sqrt{N^2 - i\omega Pe}$ and $m_2 = \sqrt{M^2 + i\omega Re}$.
$$(3.7)$$

Solving equations (3.4) and (3.5) using the boundary conditions (3.6) and (3.7), we obtain

$$u_{0}(y) = -A\cosh m_{2}y + C\frac{\sinh m_{2}y}{\sinh m_{2}} + A + B\frac{\sin m_{1}y}{\sin m_{1}}$$
(3.8)

and

whe

re
$$A = \frac{\lambda(1+\lambda_1)}{m_2^2}$$
, $B = \frac{Gr(1+\lambda_1)}{(m_1^2+m_2^2)}$ and $C = (A\cosh m_2 - A - B)$.

Therefore, the fluid velocity and temperature are given as

 $\theta_0(y) = \frac{\sin m_1 y}{\cdot}$

$$u(y,t) = \left(-A\cosh m_2 y + C\frac{\sinh m_2 y}{\sinh m_2} + A + B\frac{\sin m_1 y}{\sin m_1}\right)e^{i\omega t}$$
(3.10)

d $\theta(y,t) = \frac{\sin m_1 y}{\sin m} e^{i\omega t}$ (3.11)

RESULTS AND DISCUSSION

Fig. 2 shows the effect of material parameter λ_1 on velocity u for Re = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and M = 1. It is observed that, the axial velocity u increases with increasing λ_1 . Also, the maximum velocity occurs at the centerline of the channel while the minimum at the channel walls. Moreover, the velocity is more of Jeffrey fluid ($\lambda_1 > 0$) than that of Newtonian fluid ($\lambda_1 \to 0$).

Effect of Hartmann number M on velocity u for Re = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$ is shown in Fig.3. It is found that, the axial velocity u decreases with increasing M.

Fig. 4 depicts the effect of radiation parameter N on velocity u for Re = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, M = 1 and $\lambda_1 = 0.3$. It is noted that, the axial velocity u increases with an increase in N.

Effect of Grashof number Gr on velocity u for Re = 1, M = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$ is depicted in Fig.5. It is observed that, the axial velocity u increases with increasing Gr.

Fig. 6 illustrates the effect of Reynolds number Re on velocity u for M = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$. It is found that, the axial velocity u decreases with decreasing Re.

Effect of λ on velocity u for Re = 1, Gr = 1, Pe = 0.71, M = 1, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$ is illustrated in Fig. 7. It is noted that, the axial velocity u increases with an increase in λ .

Fig. 8 shows the effect of ω on velocity u for Re = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, M = 1, t = 0.5, N = 1 and $\lambda_1 = 0.3$. It is observed that, the axial velocity u decreases with decreasing ω .

Pelagia Research Library

2315

(3.9)

From Table-1, it is found that the axial velocity u increases with increasing Pe.

Effect of N on temperature θ for Pe = 0.71, $\omega = 1$ and t = 0.5 is shown in Fig.9. It is noted that, the temperature θ increases with an increase in N.

Fig. 10 depicts the effect of Pe on temperature θ for N=1, $\omega=1$ and t=0.5. It is observed that, the temperature θ increases with increasing Pe.

Effect of ω on temperature θ for N = 1, Pe = 0.71 and t = 0.5 is depicted in Fig. 11. It is found that, the temperature θ decreases with an increase in ω .

CONCLUSION

In this chapter, we studied the effect of heat transfer on MHD oscillatory flow of Jeffrey fluid in a channel. The expressions for the velocity and temperature are obtained analytically. It is found that, the velocity u increases with increasing λ_1, N, Gr, Pe and λ , while it decreases with increasing M, Re and ω . Also, it is observed that the temperature θ increases with increasing N and Pe, while it decreases with increasing ω . Further, it is found that, the velocity is more for Jeffrey fluid than that of Newtonian fluid.



Fig. 2 Effect of material parameter λ_1 on velocity u for Re = 1, Gr = 1,

Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and M = 1.

Pelagia Research Library



Fig. 3 Effect of Hartmann number M on velocity u for Re = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$.



Fig. 4 Effect of radiation parameter N on velocity u for Re = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, M = 1 and $\lambda_1 = 0.3$.



Fig. 5 Effect of Grash of number Gr on velocity u for Re = 1, M = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$.



Fig. 6 Effect of Reynolds number Re on velocity u for M = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$.



Fig. 7 Effect of λ on velocity u for Re = 1, Gr = 1, Pe = 0.71, M = 1, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$.



Fig. 8 Effect of ω on velocity u for Re = 1, Gr = 1, Pe = 0.71, $\lambda = 1$, M = 1, t = 0.5, N = 1 and $\lambda_1 = 0.3$.

Table-1: The variation of velocity u with y for different values of Pe with Re = 1, Gr = 1, M = 1, $\lambda = 1$, $\omega = 1$, t = 0.5, N = 1 and $\lambda_1 = 0.3$.

	Pe=0	Pe=0.5	Pe=1	Pe=2
0	0	0	0	0
0.1	0.0675	0.0679	0.0682	0.0683
0.2	0.1227	0.1236	0.1241	0.1243
0.3	0.1648	0.1660	0.1667	0.1670
0.4	0.1928	0.1942	0.1951	0.1954
0.5	0.2056	0.2070	0.2080	0.2084
0.6	0.2020	0.2035	0.2044	0.2049
0.7	0.1810	0.1822	0.1831	0.1835
0.8	0.1412	0.1421	0.1427	0.1431
0.9	0.0813	0.0818	0.0821	0.0824
1	0	0	0	0



Fig. 9 Effect of N on temperature θ for Pe = 0.71, $\omega = 1$ and t = 0.5.



Fig. 10 Effect of Pe on temperature θ for N = 1, $\omega = 1$ and t = 0.5.



Fig. 11 Effect of ω on temperature θ for N = 1, Pe = 0.71 and t = 0.5.

REFERENCES

[1] Aamir Ali and Saleem Asghar, International Journal of the Physical Sciences, 6(36)(2011), 8035 - 8043.

[2] T. K. Aldoss, M. A. Al-Nimr, M. A. Jorah, and B. Al-Shaer, Heat transfer, 28A(1995), 635-645.

[3] M.A.M. Al Khatib and S.D.R. Wilson, J. Non-Newtonian Fluid Mech., 100 (2001),1–8.

[4] H. A. Attia, and N. A. Kotb, Acta Mechanica, vol. 117 (1996), . 215-220.

[5] A.C.L. Cogley, W.G. Vinvent, and E.S. Giles, American Institute of Aeronautics and Astronautics, 6(1968), 551-553.

[6] I.A. Frigaard and D.P. Ryan, J. Non-Newtonian Fluid Mech., 123 (2004), 67-83.

[7] M. M. Hamza B.Y Isah H. Usman, International Journal of Computer Applications, 33(4) (2011), 12-17.

[8] O. D. Makinde, and P. Y. Mhone, Rom. J. Phys., 50 (2005), 931-938.

[9] A.A. Mostafa, *The Canadian journal of chemical engineering*, 87(1) (2009), 171-181.

[10] S. D. Nigam, and S. N. Singh, Q. J. Mech. Appl. Math., 13 (1960), 85-87.

[11] Raptis, C. Massias, and G. Tzivanidis, Phys. Lett., 90A(1982), 288-289.

[12] Rita and U. Jyoti das, *Physics Research International*, 2012 (2012), 1-5.

[13] V. M. Soundalgekar, and J. P. Bhat, J. Pure Appl. Math., 15 (1971), 819-828.